

Michelson interferometer operating at effects of first order with respect to v/c (the third method of measuring the speed of "aether wind")

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More than hundred years there persists a belief that Michelson-type interferometer can not be adjusted such as to detect effects of the first order with respect to v/c . Below I show that it is possible to measure the interference fringe shift (and thus the "aether wind") on the first order Michelson interferometer, and more successfully than on the interferometer of the second order. In contrast to the traditional approach, in the interferometer of the first order the light after splitting on a semi-transparent plate propagates in both arms to the reflecting mirrors in one optical medium (with dielectric permittivity ε_1), and returns after reflection from the mirrors through another optical medium (with dielectric permittivity ε_2). The shift of the interference fringe is reliably detected in the experiment when turning the interferometer by 90° . It was found to be proportional to $\varepsilon_1 - \varepsilon_2$.

Experimental data are interpreted in the bounds of the Fresnel drag of light by a moving optical medium neglecting terms quadratic in v/c . The horizontal projection v of the Earth's velocity relative to luminiferous aether thus found lies in the range $140 \text{ km/s} < v < 480 \text{ km/s}$ depending on the time of the day and night at the latitude of Obninsk. This is the third method of measuring the speed of aether wind. It gives the same range of values as two earlier described methods operating at second order with respect to v/c .

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1. TWO VARIANTS OF AETHER WIND'S DETECTOR

As is known [1, 2] there is no fringe shift in evacuated Michelson interferometer. Thus, this mode of the device is not fit to detect "aether wind". Michelson interferometer becomes sensitive to "aether wind" when using in it as light's carrier an optical medium with refractive index $n > 1$ [1, 2]. If light passes always in a single medium the interference fringe shift X_m appears to be proportional to square of the velocity v of the interferometer relative to aether: $X_m \sim (v/c)^2$, where $v \ll c$. Here the allowance for Lorentz contraction of the interferometer's longitudinal arm is vital.

Below I report results of the experiment on Michelson interferometer configured in such a way that light runs consecutively via two different optical media – it goes at first through one medium and then returns back through other medium. In this case the fringe shift proves to be proportional to v/c where $v \ll c$. So, the two-media-mode device is more sensible to "aether wind". Besides, in the interferometer of the first order the contribution of relativistic factor into the fringe shift, being quadratic with respect to v/c , can be neglected as small quantity.

2. TWO-MEDIA DETECTOR OF AETHER WIND

The device in question is a common rotary cross-like interferometer (Fig.1). Light firstly splits by the half-transparent plate into two orthogonal beams. Further each beam goes to rebounding mirrors through the optical medium having dielectric permittivity ε_1 and then returns in the medium having dielectric permittivity ε_2 . Media are separated in space by minor displacement of return containers. The respective shear of light's beams is secured by means of two mirrors. Thus, returning rays meet in the point laying at another semi-transparent plate that is angled at 45° just as the splitting plate. The space separation from the incoming beam of the light outgoing to the interference screen (1 on Fig.1) brings down parasitic noise as compared with usual construction of the second order interferometer. Next, collected on the plate 1 (Fig.1) rays are projected by the telescopic objective 2 on the screen of the vidicon 3 and then transmitted by the TV camera on the screen of the kinescope 7.

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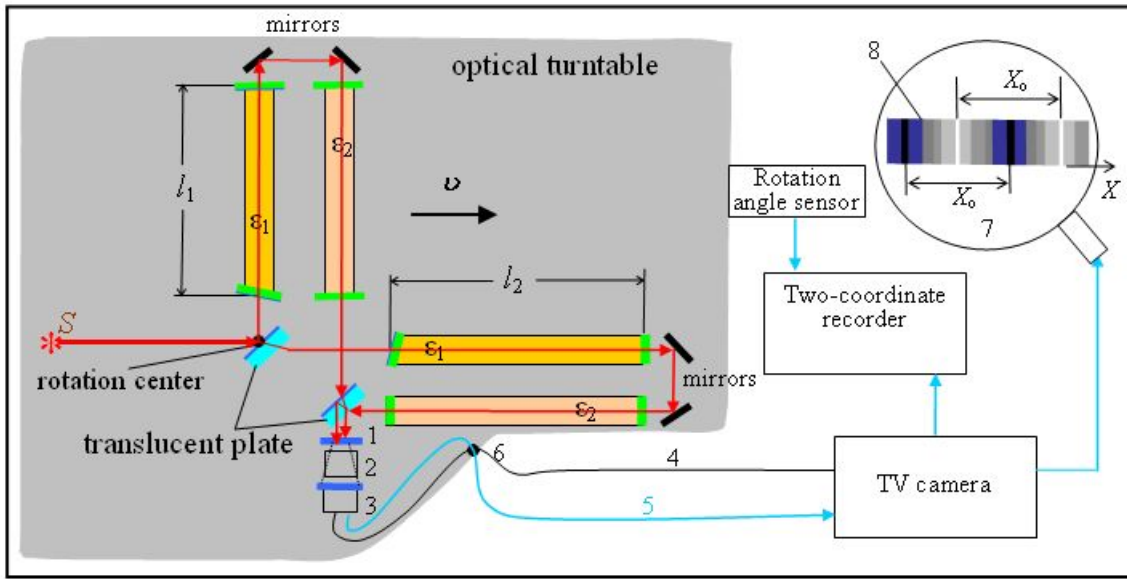


FIG. 1: Functional scheme of the first order by v/c interferometer (1971 year). Light issuing from the source S splits at the semi-transparent plate and then goes consecutively through two glass tubes filled with light's carriers of respective dielectric permittivity ε_1 and ε_2 . Ends of the tubes are covered by thin glass lids. After reflecting from rebounding mirrors the light returns for interference at the other half-transparent plate. 1 is the interference screen, 2 telescopic ocular, 3 vidicon with a deflecting facility, 4 and 5 are feeding and video cords passing via the tube 6 in the center of rotation, 8 is the interference pattern at the screen of the kinescope 7.

3. COMPUTING AETHER WIND VELOCITY FROM INTERFERENCE FRINGE SHIFT

The interference fringe shift is proportional to difference Δt of round-trip times according to

$$X_m = c \frac{X_o}{\lambda} \Delta t \quad (1)$$

where X_o is the bandwidth and λ the wavelength. By virtue of (1) we may reason in terms of time intervals t .

The speed of light \tilde{c} in the moving with velocity $v > 0$ optical medium is given by Fresnel formula

$$\tilde{c}'_{\pm} = \frac{c}{n} \pm v \left(1 - \frac{1}{n^2}\right) = \frac{c}{\sqrt{\varepsilon}} \pm v \frac{\Delta \varepsilon}{\varepsilon} \quad (2)$$

as measured in the reference frame of the moving medium. Here n is the refractive index and $\Delta \varepsilon = n^2 - 1$ the contribution of particles into the dielectric permittivity $\varepsilon = n^2$ of the optical medium.

From (2) we have for the round-trip time t_{\parallel} in the direction of interferometer arm parallel to v

$$t_{\parallel} = \frac{l'}{\tilde{c}'_+} + \frac{l'}{\tilde{c}'_-} = \frac{l'}{c/\sqrt{\varepsilon_1} + v\Delta\varepsilon_1/\varepsilon_1} + \frac{l'}{c/\sqrt{\varepsilon_2} - v\Delta\varepsilon_2/\varepsilon_2} \approx \frac{l}{c} \left[\sqrt{\varepsilon_1} + \sqrt{\varepsilon_2} + \frac{v}{c} (\Delta\varepsilon_1 - \Delta\varepsilon_2) \right] \quad (3)$$

where ε_1 is the dielectric permittivity of the forward medium, and ε_2 that of the return medium. In deriving (3) terms with $(v/c)^2$ were discarded.

We have for the round-trip time in the orthogonal direction

$$t_{\perp} = \frac{l}{\sqrt{(c/n_1)^2 + v^2}} + \frac{l}{\sqrt{(c/n_2)^2 + v^2}} \approx \frac{l}{c} (\sqrt{\varepsilon_1} + \sqrt{\varepsilon_2}) \quad (4)$$

where again terms with $(v/c)^2$ were neglected.

Subtracting (3) from (4) gives

$$\Delta t = t_{\perp} - t_{\parallel} \approx \frac{v l}{c c} (\Delta\varepsilon_1 - \Delta\varepsilon_2) = \frac{v l}{c c} (\varepsilon_1 - \varepsilon_2). \quad (5)$$

Formula (5) shows that the difference of round-trip times depends on the difference of the contribution of particles into dielectric permittivities of two light carrying media.

4. MEASURING INTERFERENCE FRINGE SHIFT

Fig.2 shows the measured amplitude A_m of the harmonic component of the interference fringe shift as a function of local time t_{local} at the attitude of Obninsk. Measurements covered 24 hours of the day and night. The light's carriers were: in the forward direction carbon bisulfide (CS_2) $\Delta\varepsilon_1 = 0.0036$, and on the return path the air at normal pressure, $\Delta\varepsilon_1 = 0.0006$. The right scale of ordinates shows values of the aether wind velocity $v_{hor.} \sim A_m$ km/s. The dependence $v(t_{local})$ agrees well with that obtained on the interferometer of second order with respect to v/c [1, 2].

According to (5) the first order device is $(v/c)^{-1} \approx 1000$ times more sensitive than the interferometer of the second order with respect to v/c . This enabled me to use a relatively small arms $l = 0.2$ m.

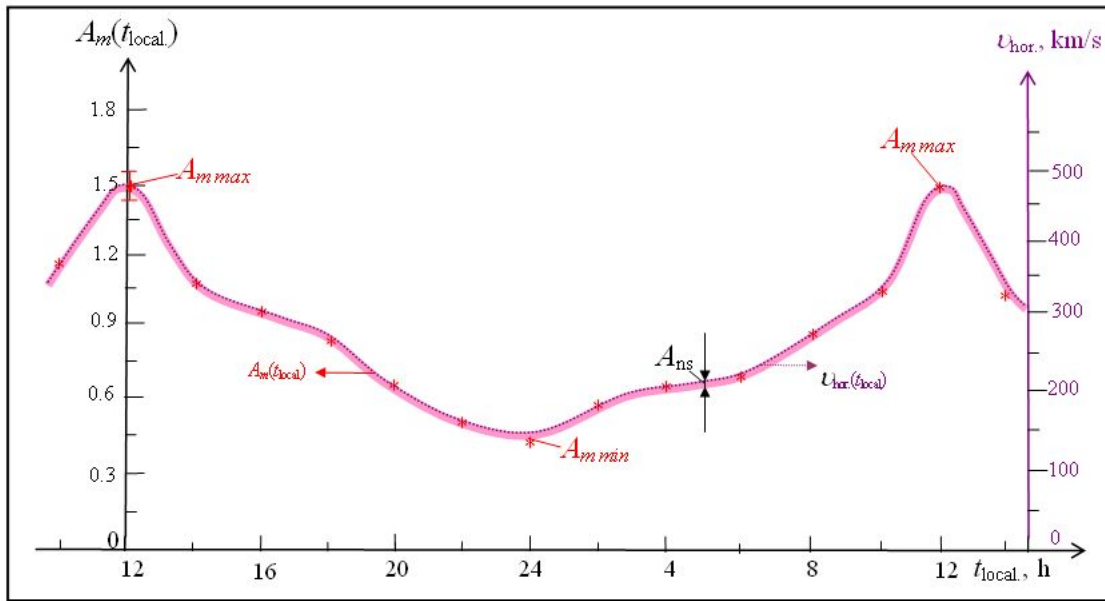


FIG. 2: Measured relative amplitude $A_m = X_m/X_o$ at various times of day and night, local time at Obninsk 55.8° NL, 22 June 1971 year, where X_m is the amplitude of interference fringe shift and $X_o = 90$ mm is the bandwidth on the kinescope's screen. Light's carriers are CS_2 having $\varepsilon \approx 0.0036$ and air having $\varepsilon \approx 0.0006$. The line width shows the jitter noise A_{ns} of the interferometer. The horizontal projection $v_{hor.}$ of the aether wind velocity, computed from A_m by (5) with (1) is shown at the right ordinate axis. Parameters of the experiment: arms are $l_1 = l_2 = 0.2$ m, wavelength $\lambda = 6 \cdot 10^{-7}$ m.

Fig.3 shows the dependence of the interference fringe shift X_m on the difference $\varepsilon_1 - \varepsilon_2$ between dielectric permittivities of light's carriers. As we see, X_m linearly grows from $\varepsilon_1 - \varepsilon_2 = 0$ to $\varepsilon_1 - \varepsilon_2 = 1$. By (5) the sensitivity can be enhanced in 10^5 times (100 times due to $\varepsilon_1 - \varepsilon_2$ and 1000 times due to v/c). So that the measuring can be performed at microwaves, i.e. 10^5 times greater wavelengths than in optics. The microwave experiment was made using the $CaTiO_3$ ferroelectric guide, $\varepsilon = 255$, at $\lambda = 10$ cm by the direct measurement of the phase difference. The converted to X_m magnitude lies exactly on the continuation of the line 1 in Fig.3 [1].

5. CONCLUSION

Thus the reality of the first order with respect to v/c Michelson interferometer is experimentally demonstrated. Experiments show that the shift of the interference fringe occurs in optical media and is absent in vacuum. The first order interferometer is $(v/c)^{-1} \approx 1000$ times more sensitive to aether wind than the second order Michelson interferometer. The first order interferometer shows the ratio signal/noise ~ 100 all round the clock. Such the resolving power and stability of measurements at any time of day and night and any season attests confidently the presence of the absolute reference frame with the Earth moving relative to it with the velocity that is above 480 km/s.

Earlier the same estimation of the "aether wind" velocity has been obtained on Michelson interferometer by two methods [3] operating at second order effects with respect to v/c [2]. It was the historically first method that took into account parameters of optical medium specially used as a light's carrier. The content presented here is actually a third method of experimental measuring the velocity of "aether wind" via the shift of interference fringe at the

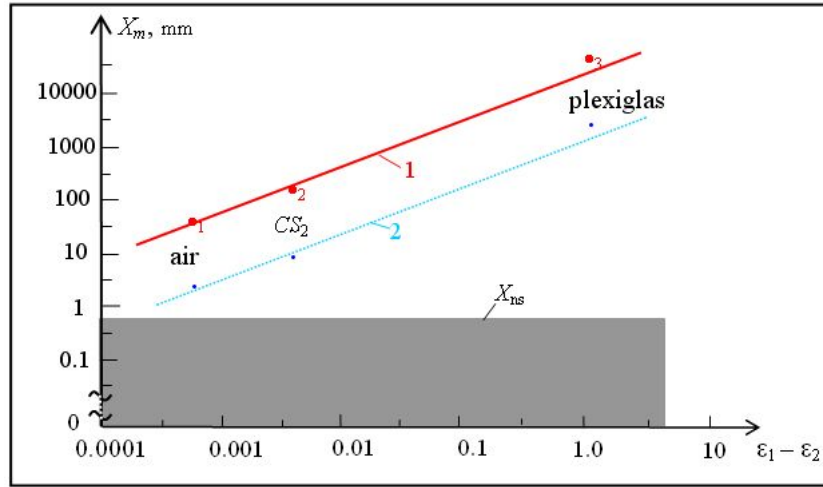


FIG. 3: Measured shift X_m of the interference fringe as a function of difference $\epsilon_1 - \epsilon_2$ between dielectric permittivities of two light's carriers: ●₁ air ($\epsilon_1 = 1.0006$)/lab vacuum ($\epsilon_2 = 1.000006$), ●₂ CS₂ ($\epsilon_1 = 1.0036$)/air, ●₃ plexiglas ($\epsilon_1 \approx 2$)/air. Abscissa and ordinate scales are logarithmic. Line 1 corresponds to $X_{m \max}$, line 2 to $X_{m \min}$ in notations of Fig.2. δX_{ns} is the jitter noise read on the screen of kinescope. Parameters of the experiment for CS₂: arms are $l_1 = l_2 = 0.2$ m, wavelength $\lambda = 6 \cdot 10^{-7}$ m (all other data are reduced to this values)

transverse beams interferometer – the method operating at first order effects with respect to v/c due to use of two optical media as light's carriers.

Thus, kinetic evidences for existence of luminiferous aether are to this time provided by three phase interferometry techniques, two operating at second order effects and third one, described in the present report, at the first order one.

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